# Unit vectors and metric coefficients

Spherical coordinates  $r, \theta, \phi$  are defined by

$$x = r \sin \theta \cos \phi$$
  

$$y = r \sin \theta \sin \phi$$
 (1)  

$$z = r \cos \theta$$

They are the coordinates of choice in problems with spherical boundaries. Since z is no longer one of the coordinates we will not be able to use  $A_z$  and  $F_z$  to specify the fields. This will lead to some subtleties. Let's first derive the unit vectors and the metric coefficients. Since

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Vectors and operators in spherical

coordinates

$$\frac{\partial \mathbf{r}}{\partial r} = (\sin\theta\cos\phi, \sin\theta\sin\phi, \cos\theta)$$
(2)

we have

$$h_r = \left| \frac{\partial \mathbf{r}}{\partial r} \right| = \sqrt{\sin^2 \theta [\cos^2 \phi + \sin^2 \phi] + \cos^2 \theta} = 1$$
(3)

and

$$\hat{a}_r = (\sin\theta\cos\phi, \sin\theta\sin\phi, \cos\theta)$$
 (4)

Likewise

$$\frac{\partial \mathbf{r}}{\partial \theta} = r \left( \cos \theta \cos \phi, \cos \theta \sin \phi, -\sin \theta \right)$$
(5)

so that

$$h_{\theta} = \left| \frac{\partial \mathbf{r}}{\partial \theta} \right| = r \sqrt{\cos^2 \theta \left[ \cos^2 \phi + \sin^2 \phi \right] + \sin^2 \theta} = r \tag{6}$$

and

$$\hat{a}_{\theta} = (\cos\theta\cos\phi, \cos\theta\sin\phi, -\sin\theta)$$
(7)

Finally

$$\frac{\partial \mathbf{r}}{\partial \phi} = r(-\sin\theta\sin\phi, \sin\theta\cos\phi, 0) \tag{8}$$

so

$$h_{\phi} = \left| \frac{\partial \mathbf{r}}{\partial \phi} \right| = r \sqrt{\sin^2 \theta \left[ \cos^2 \phi + \sin^2 \phi \right]} = r \sin \theta \tag{9}$$

and

$$\hat{a}_{\phi} = (-\sin\phi, \cos\phi, 0) \tag{10}$$

Using the metric coefficients, we are now ready to derive the various differential operators.

## **Differential operators**

Using the results of Lecture 1b for a general orthogonal coordinate system, and identifying  $u=r, v=\theta, w=\phi$  we have the gradient

$$\nabla f = \hat{a}_r \frac{\partial f}{\partial r} + \hat{a}_{\theta} \frac{1}{r} \frac{\partial f}{\partial \theta} + \hat{a}_{\phi} \frac{1}{r \sin \theta} \frac{\partial f}{\partial \phi}$$
(11)

The divergence is

$$\nabla \cdot \mathbf{A} = \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial r} (r^2 \sin \theta A_r) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} (r \sin \theta A_\theta) + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \phi} (r A_\phi)$$
(12)

Simplifying we obtain

$$\nabla \cdot \mathbf{A} = \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 A_r \right) + \frac{1}{r \sin \theta} \left[ \frac{\partial}{\partial \theta} \left( \sin \theta A_\theta \right) + \frac{\partial}{\partial \phi} A_\phi \right]$$
(13)

The curl is

$$\nabla \times \mathbf{A} = \frac{\hat{a}_{r}}{r^{2} \sin \theta} \left[ \frac{\partial}{\partial \theta} (r \sin \theta A_{\phi}) - \frac{\partial}{\partial \phi} (r A_{\theta}) \right] + \frac{\hat{a}_{\theta}}{r \sin \theta} \left[ \frac{\partial}{\partial \phi} A_{r} - \frac{\partial}{\partial r} (r \sin \theta A_{\phi}) \right]$$
(14)
$$+ \frac{\hat{a}_{\phi}}{r} \left[ \frac{\partial}{\partial r} (r A_{\theta}) - \frac{\partial}{\partial \theta} A_{r} \right]$$

Simplified (somewhat) we get

$$\nabla \times \mathbf{A} = \frac{\hat{a}_{r}}{r \sin \theta} \left[ \frac{\partial}{\partial \theta} (\sin \theta A_{\phi}) - \frac{\partial}{\partial \phi} A_{\theta} \right] \\ + \frac{\hat{a}_{\theta}}{r} \left[ \frac{1}{\sin \theta} \frac{\partial}{\partial \phi} A_{r} - \frac{\partial}{\partial r} (r A_{\phi}) \right] \\ + \frac{\hat{a}_{\phi}}{r} \left[ \frac{\partial}{\partial r} (r A_{\theta}) - \frac{\partial}{\partial \theta} A_{r} \right]$$
(15)

The Laplacian is

$$\nabla^{2} f = \frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial r} \left( r^{2} \sin \theta \frac{\partial f}{\partial r} \right) + \frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta} \left( \frac{r \sin \theta}{r} \frac{\partial f}{\partial \theta} \right) + \frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \phi} \left( \frac{r}{r \sin \theta} \frac{\partial f}{\partial \phi} \right)$$
(16)

This simplifies to

$$\nabla^{2} f = \frac{1}{r^{2}} \frac{\partial}{\partial r} \left( r^{2} \frac{\partial f}{\partial r} \right) + \frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial f}{\partial \theta} \right)$$

$$+ \frac{1}{r^{2} \sin^{2} \theta} \frac{\partial^{2} f}{\partial \phi^{2}}$$
(17)

The Laplacian of a vector field is by definition

$$\nabla^2 \mathbf{A} \equiv \nabla (\nabla \cdot \mathbf{A}) - \nabla \times \nabla \times \mathbf{A}$$
(18)

In cylindrical coordinates we avoided this expression because for  $\mathbf{A} = \hat{a}_z A_z$  will simply had  $\nabla^2 \mathbf{A} = \hat{a}_z \nabla^2 A_z$ . Things are not as simple in spherical coordinates.

## **Helmholtz** equation

The Helmholtz equation, which is independent of any coordinate system, is

$$\nabla^2 \mathbf{A} + \beta^2 \mathbf{A} = 0 \tag{19}$$

In general this will read

$$\nabla (\nabla \cdot \mathbf{A}) - \nabla \times \nabla \times \mathbf{A} + \beta^2 \mathbf{A} = 0$$
 (20)

This is a pretty ugly vector equation. However, recall from Lecture 2c that

$$\nabla \cdot \mathbf{A} = \frac{j}{\omega} \left( \nabla \cdot \mathbf{E} + \nabla^2 \psi \right)$$
(21)

where  $\Psi$  is an arbitrary scalar field. The divergence of **A** is completely unconstrained by physics. We can make it be anything we want. Let's see if we can use this freedom to reduce our vector equation to a single scalar equation. Let's take

$$\mathbf{A} = \hat{a}_r A_r \tag{22}$$

From the curl expression we see that these will be  $TM^r$ modes in which  $H_r \equiv 0$ . We have

$$\nabla \times \mathbf{A} = \frac{\hat{a}_{\theta}}{r} \frac{1}{\sin \theta} \frac{\partial}{\partial \phi} A_r - \frac{\hat{a}_{\phi}}{r} \frac{\partial}{\partial \theta} A_r \qquad (23)$$

Applying another curl we obtain

$$\nabla \times \nabla \times \mathbf{A} = \frac{\hat{a}_r}{r\sin\theta} \left[ -\frac{\partial}{\partial\theta} \left( \sin\theta \frac{1}{r} \frac{\partial}{\partial\theta} A_r \right) - \left( \frac{\partial}{\partial\phi} \frac{1}{r\sin\theta} \frac{\partial}{\partial\phi} A_r \right) \right] \quad (24)$$
$$+ \frac{\hat{a}_{\theta}}{r} \frac{\partial}{\partial r} \left( r \frac{1}{r\partial\theta} A_r \right) + \frac{\hat{a}_{\phi}}{r} \frac{\partial}{\partial r} \left( r \frac{1}{r\sin\theta} \frac{\partial}{\partial\phi} A_r \right)$$

Therefore

$$\nabla \times \nabla \times \mathbf{A} = \frac{\hat{a}_r}{r^2 \sin \theta} \left[ \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} A_r \right) + \frac{1}{\sin \theta} \left( \frac{\partial^2}{\partial^2 \phi^2} A_r \right) \right]$$
(25)  
$$- \frac{\hat{a}_{\theta}}{r} \frac{\partial}{\partial r} \frac{\partial}{\partial \theta} A_r - \frac{\hat{a}_{\phi}}{r \sin \theta} \frac{\partial}{\partial r} \frac{\partial}{\partial \phi} A_r$$

Since  $\beta^2 \mathbf{A}$  has only an *r* component, let's see if we can use  $\nabla \cdot \mathbf{A}$  to cancel the  $\theta, \phi$  components of  $-\nabla \times \nabla \times \mathbf{A}$ . Let's

call  $\nabla \cdot \mathbf{A} = f$  and remember that this scalar function can be anything we want it to be.  $\nabla (\nabla \cdot \mathbf{A})$  is

$$\nabla f = \hat{a}_r \frac{\partial f}{\partial r} + \hat{a}_\theta \frac{1}{r} \frac{\partial f}{\partial \theta} + \hat{a}_\phi \frac{1}{r \sin \theta} \frac{\partial f}{\partial \phi}$$
(26)

If

$$\frac{1}{r}\frac{\partial f}{\partial \theta} - \frac{1}{r}\frac{\partial}{\partial r}\frac{\partial}{\partial \theta}A_r = 0$$
(27)

and

$$\frac{1}{r\sin\theta}\frac{\partial f}{\partial\phi} - \frac{1}{r\sin\theta}\frac{\partial}{\partial r}\frac{\partial}{\partial\phi}A_r = 0$$
(28)

then we will have achieved our goal. This is true if we take

$$f = \frac{\partial}{\partial r} A_r \tag{29}$$

The operator  $\nabla(\nabla \cdot \mathbf{A}) - \nabla \times \nabla \times \mathbf{A}$  now has only the *r* component

$$\frac{\partial^2 A_r}{\partial r^2} + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial A_r}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 A_r}{\partial \phi^2} \qquad (30)$$

where  $\partial^2 A_r / \partial r^2$  is the *r* component of  $\nabla f$ . We could add  $\beta^2 A_r$  to this and apply separation of variables. However, from a mathematical point of view it would be much preferred to if we could employ the scalar Laplacian operator. Now  $\nabla^2 A_r$  is

$$\frac{1}{r^2}\frac{\partial}{\partial r}\left(r^2\frac{\partial A_r}{\partial r}\right) + \frac{1}{r^2\sin\theta}\frac{\partial}{\partial\theta}\left(\sin\theta\frac{\partial A_r}{\partial\theta}\right) + \frac{1}{r^2\sin^2\theta}\frac{\partial^2 A_r}{\partial\phi^2} \quad (31)$$

This is almost the same as (30), but the first terms differ. In fact

$$\frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial A_r}{\partial r} \right) = \frac{\partial^2 A_r}{\partial r^2} + \frac{2}{r} \frac{\partial A_r}{\partial r}$$
(32)

and we have an extra  $(2/r)\partial A_r/\partial r$  term. However, note that

$$\frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial}{\partial r} \left( \frac{A_r}{r} \right) \right) = \frac{1}{r} \frac{\partial^2 A_r}{\partial r^2}$$
(33)

Therefore

$$\nabla^2 \left(\frac{A_r}{r}\right) = \frac{1}{r} \left[ \frac{\partial^2 A_r}{\partial r^2} + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial A_r}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 A_r}{\partial \phi^2} \right]$$
(34)

This tells us that the function  $\psi = A_r/r$  satisfies the scalar Helmholtz equation

$$\nabla^2 \psi + \beta^2 \psi = 0 \tag{35}$$

in spherical coordinates. If we solve this equation, then the vector potential is obtained from  $A_r = r \psi$ .

## **Separation of variables**

Let's take

$$\psi = f(r)g(\theta)h(\phi) \tag{36}$$

The Helmholtz equation then reads

$$\frac{1}{r^{2}}\frac{\partial}{\partial r}(r^{2}f'gh) + \frac{1}{r^{2}\sin\theta}\frac{\partial}{\partial\theta}(\sin\theta fg'h) + \frac{1}{r^{2}\sin^{2}\theta}fgh'' + \beta^{2}fgh = 0$$
(37)

Dividing through by *fgh* we obtain

$$\frac{1}{r^{2}f}\frac{\partial}{\partial r}(r^{2}f') + \frac{1}{r^{2}g\sin\theta}\frac{\partial}{\partial\theta}(\sin\theta g') + \frac{1}{r^{2}\sin^{2}\theta}\frac{h''}{h} + \beta^{2} = 0$$
(38)

Multiplying through by  $r^2 \sin^2 \theta$  we arrive at

$$\frac{\sin^{2}\theta}{f} \frac{\partial}{\partial r} (r^{2} f') + \frac{\sin\theta}{g} \frac{\partial}{\partial \theta} (\sin\theta g') + \frac{h''}{h} + \beta^{2} r^{2} \sin^{2}\theta = 0$$
(39)

The h''/h term depends only on  $\phi$  and no other term depends on  $\phi$ . By our usual argument this term must be a constant. Let's write

$$\frac{h^{\prime\prime}}{h} = -m^2 \tag{40}$$

where  $m^2$  is an arbitrary complex number (but *m* will typically be an integer) and

$$h = \begin{cases} \cos(m\,\phi) \\ \sin(m\,\phi) \end{cases} \tag{41}$$

Replacing h''/h in the Helmholtz equation and dividing through by  $\sin^2 \theta$  we arrive at

$$\frac{1}{f}\frac{\partial}{\partial r}(r^2 f') + \beta^2 r^2 + \frac{1}{g\sin\theta}\frac{\partial}{\partial \theta}(\sin\theta g') - \frac{m^2}{\sin^2\theta} = 0 \quad (42)$$

Notice that all *r* dependence is in the first two terms and all  $\theta$  dependence is in the last two terms. Therefore

$$\frac{1}{g\sin\theta} \frac{\partial}{\partial\theta} (\sin\theta g') - \frac{m^2}{\sin^2\theta} = -n(n+1)$$
(43)

where -n(n+1) is an arbitrary complex constant (although *n* will usually be an integer). We have

$$\frac{1}{\sin\theta} \frac{\partial}{\partial\theta} \left( \sin\theta g' \right) + \left[ n(n+1) - \frac{m^2}{\sin^2\theta} \right] g = 0$$
(44)

-

or

$$g'' + \frac{\cos\theta}{\sin\theta}g' + \left[n(n+1) - \frac{m^2}{\sin^2\theta}\right]g = 0$$
(45)

We will come back to this equation later. What remains of the Helmholtz equation is the r dependence

$$\frac{\partial}{\partial r} \left( r^2 f' \right) + \left[ \beta^2 r^2 - n(n+1) \right] f = 0$$
(46)

which expands to

$$r^{2} f'' + 2r f' + \left[\beta^{2} r^{2} - n(n+1)\right] f = 0$$
(47)

Let's call

$$\begin{aligned} x &= \beta r \\ y(x) &= f(r) \end{aligned}$$
 (48)

Using the Chain Rule we have

$$\frac{df}{dr} = \frac{dy}{dx}\frac{dx}{dr} = \beta y'$$
(49)

and our equation becomes

$$x^{2}y'' + 2xy' + [x^{2} - n(n+1)]y = 0$$
(50)

If it wasn't for the factor of 2 this would be in the form of Bessel's equation

$$x^{2} y'' + x y' + (x^{2} - v^{2}) y = 0$$
 (51)

Since we've already solved Bessel's equation, it makes sense to look for a transformation that will covert our equation to Bessel's equation. Setting  $y(x)=z(x)/\sqrt{x}$  we have

$$y' = \frac{z'}{\sqrt{x}} - \frac{z}{2x^{3/2}}$$
  
$$y'' = \frac{z''}{\sqrt{x}} - \frac{z'}{x^{3/2}} + \frac{3}{4} \frac{z}{x^{5/2}}$$
 (52)

Substituting these expressions into (50) and multiplying through by  $\sqrt{x}$  results in

$$x^{2}z'' + xz' + [x^{2} - (n+1/2)^{2}]z = 0$$
(53)

which is Bessel's equation with v=n+1/2. The solutions are linear combinations of the Bessel functions

$$z = \begin{cases} J_{n+1/2}(x) \\ Y_{n+1/2}(x) \end{cases}$$
(54)

Therefore y(x) is  $1/\sqrt{x}$  times a Bessel function. We are led to define the *spherical Bessel functions* 

$$j_{n}(x) = \sqrt{\frac{\pi}{2x}} J_{n+1/2}(x)$$

$$y_{n}(x) = \sqrt{\frac{\pi}{2x}} Y_{n+1/2}(x)$$
(55)

The *r* behavior of  $\Psi$  is therefore a linear combination of  $j_n(\beta r)$  and  $y_n(\beta r)$ . Since  $A_r = r \Psi$ , our solutions will have

the form

$$A_{r} = r \begin{cases} j_{n}(\beta r) \\ y_{n}(\beta_{r}) \end{cases} \begin{cases} P_{n}^{m}(\cos \theta) \\ Q_{n}^{m}(\cos \theta) \end{cases} \begin{cases} \cos(m \phi) \\ \sin(m \phi) \end{cases}$$
(56)

where  $P_n^m(\cos\theta)$ ,  $Q_n^m(\cos\theta)$  are Legendre functions. We will derive these in the next lecture.

# Fields

#### TM<sup>r</sup> modes

Using (15) and  $\mathbf{H} = (1/\mu)\nabla \times \mathbf{A}$ ,  $\mathbf{E} = (-j/\omega \epsilon)\nabla \times \mathbf{H}$  with  $\mathbf{A} = \hat{a}_r A_r$ , we have

$$H_{r}=0$$

$$H_{\theta}=\frac{1}{\mu r \sin \theta} \frac{\partial}{\partial \phi} A_{r}$$

$$H_{\phi}=-\frac{1}{\mu r} \frac{\partial}{\partial \theta} A_{r}$$

$$E_{r}=\frac{j}{\omega \mu \epsilon r^{2} \sin \theta} \left[ \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} A_{r} \right) + \frac{1}{\sin \theta} \frac{\partial^{2}}{\partial \phi^{2}} A_{r} \right]$$

$$E_{\theta}=-\frac{j}{\omega \mu \epsilon r} \frac{\partial}{\partial r} \frac{\partial}{\partial \theta} A_{r}$$

$$E_{\phi}=-\frac{j}{\omega \mu \epsilon r \sin \theta} \frac{\partial}{\partial r} \frac{\partial}{\partial \phi} A_{r}$$
(57)

 $TE^r$  modes

Using (15) and  $\mathbf{E} = -(1/\epsilon)\nabla \times \mathbf{F}$ ,  $\mathbf{H} = (j/\omega\mu)\nabla \times \mathbf{E}$  with  $\mathbf{F} = \hat{a}_r F_r$ , we have

$$E_{r}=0$$

$$E_{\theta}=-\frac{1}{\epsilon r \sin \theta} \frac{\partial}{\partial \phi} F_{r}$$

$$E_{\phi}=\frac{1}{\epsilon r} \frac{\partial}{\partial \theta} F_{r}$$

$$H_{r}=\frac{j}{\omega \mu \epsilon r^{2} \sin \theta} \left[ \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} F_{r} \right) + \frac{1}{\sin \theta} \frac{\partial^{2}}{\partial \phi^{2}} F_{r} \right]$$

$$H_{\theta}=-\frac{j}{\omega \mu \epsilon r} \frac{\partial}{\partial r} \frac{\partial}{\partial \theta} F_{r}$$

$$H_{\phi}=-\frac{j}{\omega \mu \epsilon r \sin \theta} \frac{\partial}{\partial r} \frac{\partial}{\partial \phi} F_{r}$$
(58)

#### Appendix

The following Maxima code generates the spherical Bessel functions  $j_n(x)$  and their Taylor series.

N:3\$
j[0]:sin(x)/x\$
j[1]:sin(x)/x^2-cos(x)/x\$
for n:1 step 1 thru N-1 do (
 j[n+1]:ratsimp(((2\*n+1)/x)\*j[n]-j[n-1])
)\$
for n:0 step 1 thru N do (
 display(j[n]),
 j[n]:taylor(j[n],x,0,n+2),
 display(j[n])
)\$

The output is (notation  $j_n = j_n(x)$ )

$$j_{0} = \frac{\sin x}{x}$$

$$j_{0} = 1 - \frac{x^{2}}{6} + \cdots$$

$$j_{1} = \frac{\sin x}{x^{2}} - \frac{\cos x}{x}$$

$$j_{1} = \frac{x}{3} - \frac{x^{3}}{30} + \cdots$$

$$j_{2} = -\frac{(x^{2} - 3)\sin x + 3x\cos x}{x^{3}}$$

$$j_{2} = \frac{x^{2}}{15} - \frac{x^{4}}{210} + \cdots$$

$$j_{3} = -\frac{(6x^{2} - 15)\sin x + (15x - x^{3})\cos x}{x^{4}}$$

$$j_{3} = \frac{x^{3}}{105} - \frac{x^{5}}{1890} + \cdots$$

The Taylor series show that  $j_n(x)$  behaves as  $x^n$  for small x.